

Ex situ metrology of x-ray diffraction gratings (with interferometric microscopes)

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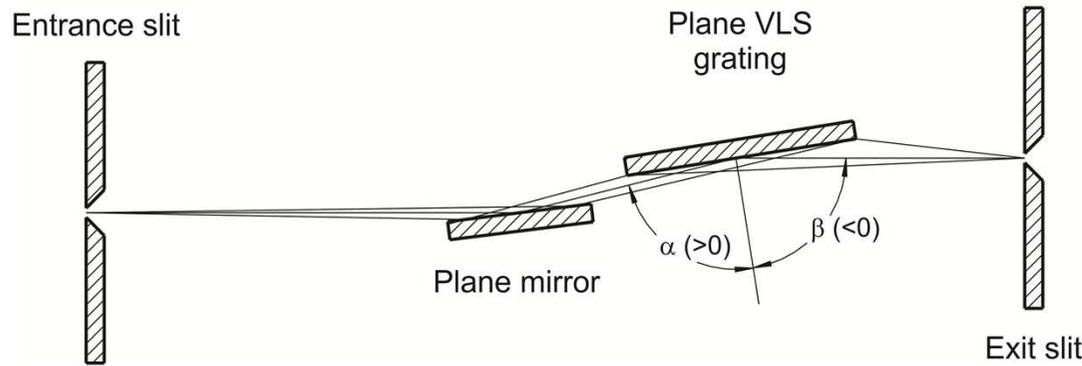


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A plane grating with a varied line spacing (VLS) in the SX700-type monochromator geometry

(The ALS MAESTRO project)



- VLS grating does all of the focusing in the modified SX700 style monochromator.

- The fixed-focus parameter

$$C = \text{Cos} \beta / \text{Cos} \alpha$$

is kept constant during the energy scan.

- Then, the ratio of virtual-source, r' , and real-source, r , distances also stays fixed.

VLS groove density: $g = g_0 + g_1 \cdot w + g_2 \cdot w^2$

Diffraction equation: $\text{Sin} \alpha + \text{Sin} \beta = m \lambda g_0$

Focusing equation of a VLS plane grating: $\frac{\text{Cos}^2 \alpha}{r} + \frac{\text{Cos}^2 \beta}{r'} = m \lambda g_1$

- **Zeroing the defocusing term:**

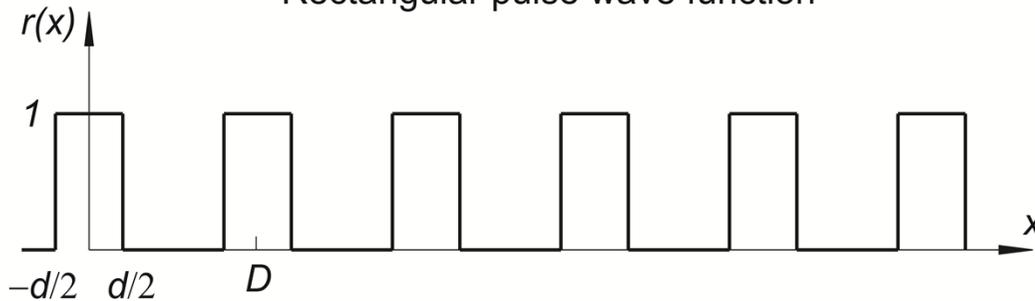
$$\text{Sin}^2 \beta \left(-\frac{1}{r} - \frac{1}{r'} \right) + \text{Sin} \beta \left(\frac{2m \lambda g_0}{r} \right) + \left[\frac{1}{r} + \frac{1}{r'} - m \lambda g_1 - \frac{(2m \lambda g_1)^2}{r} \right] = 0$$

- **The second order term is chosen (usually numerically) to minimize higher order aberrations.**

? OPTICAL METROLOGY ?

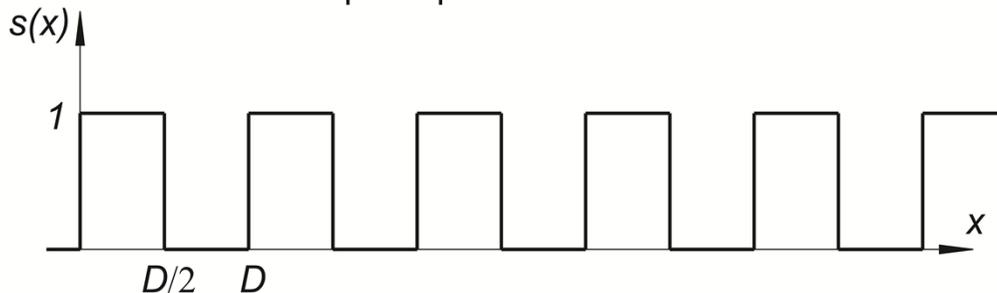
The Fourier series expansion for a rectangular pulse wave function

Rectangular pulse wave function



$$r(x) = \frac{d}{D} + \sum_{n=1}^{\infty} \frac{2}{\pi n} \sin\left(\frac{\pi n d}{D}\right) \cos\left(\frac{2\pi n}{D} x\right)$$

Square pulse wave function



$$s(x) = \frac{1}{2} + \frac{4}{\pi} \left[\sin\left(\frac{2\pi}{D} x\right) + \frac{1}{3} \sin\left(\frac{2\pi}{D} 3x\right) + \frac{1}{5} \sin\left(\frac{2\pi}{D} 5x\right) + \dots \right]$$

The first (and all other!) harmonics contains information on groove density:

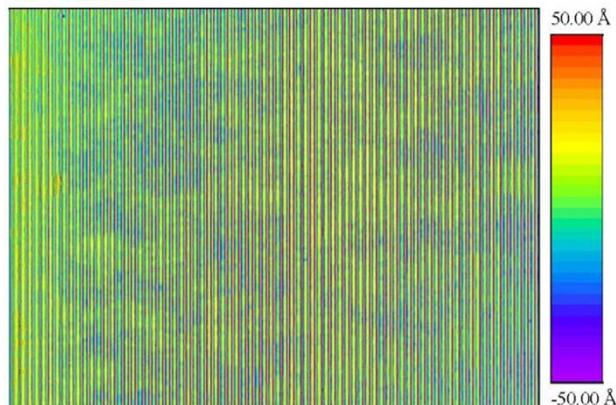
$$\frac{1}{D}$$

- **Metrology Method:** Fourier (Power Spectral Density, PSD) analysis of grating profile measurements.

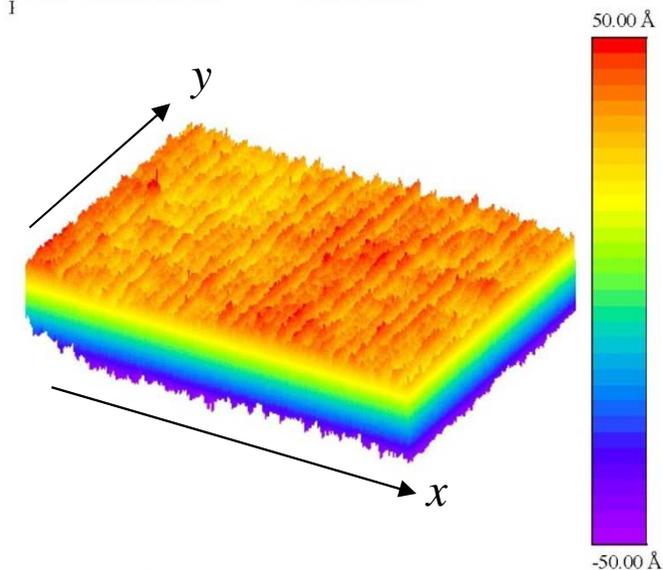
- **Limiting factors:** Grating profile measurements are performed over a limited area with a limited resolution

PSD measurements with MicroMap™-570 of X-ray grating with variable groove density

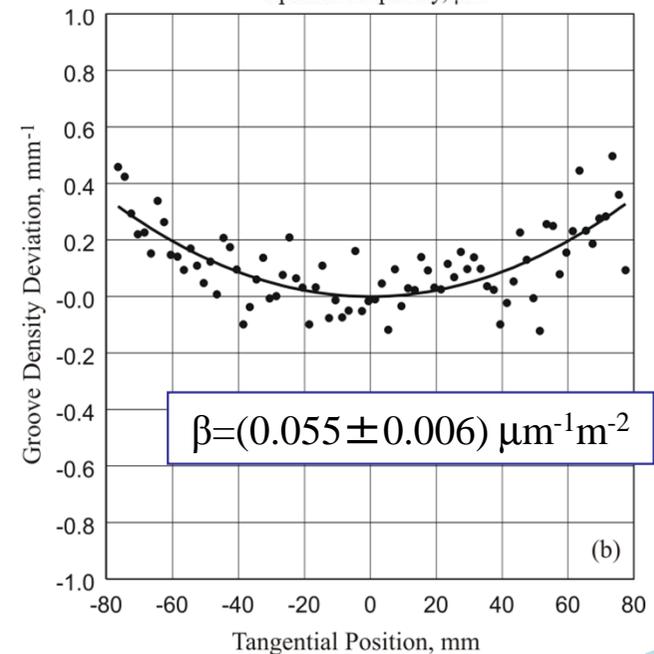
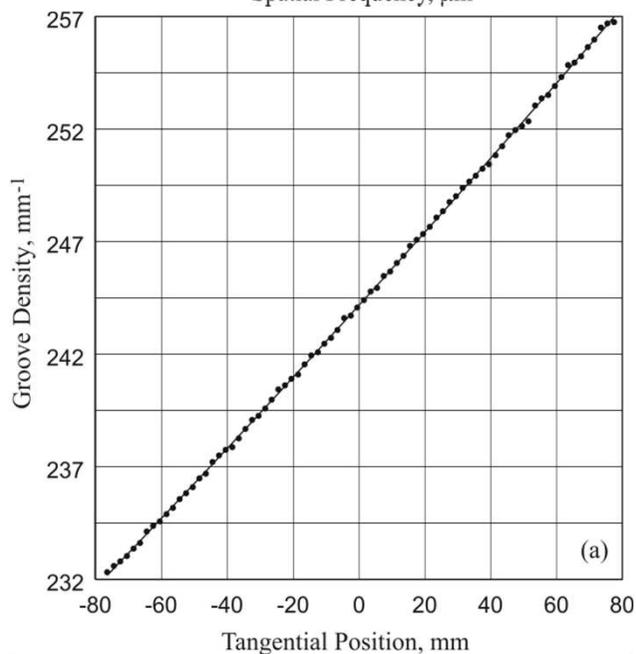
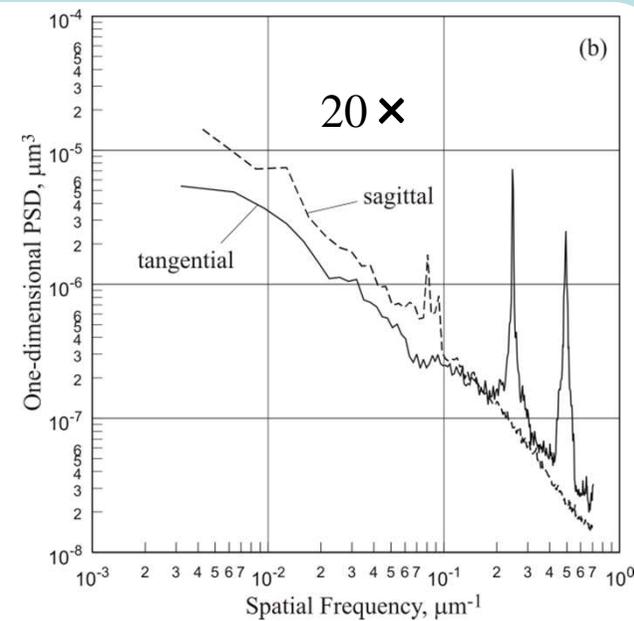
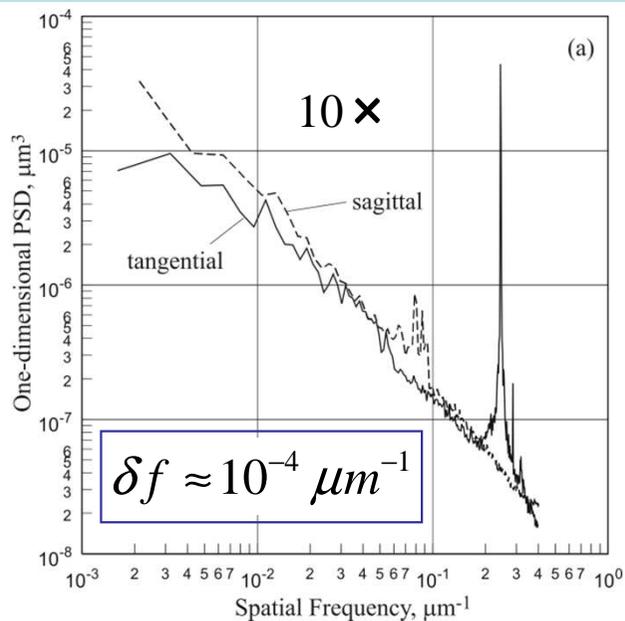
G201_Ni_x085_center_y24_10X_ PN: y=24 mm
 B6012 G201 Ni 520 mm
 2004-12-03 11:17:53 10X
 Op: vyy Points: 307200
 Lot: PtNi VLS



Sq: 28.07 Å Smooth Phase
 Sa: 24.99 Å Area: 627.2 x 470.4 μm
 St: 154.8 Å Points: 307200
 QUARTIC 10X
 I Area: 627.2 x 470.4 μm Smooth Phase
 I



Sq: 28.07 Å 640x480



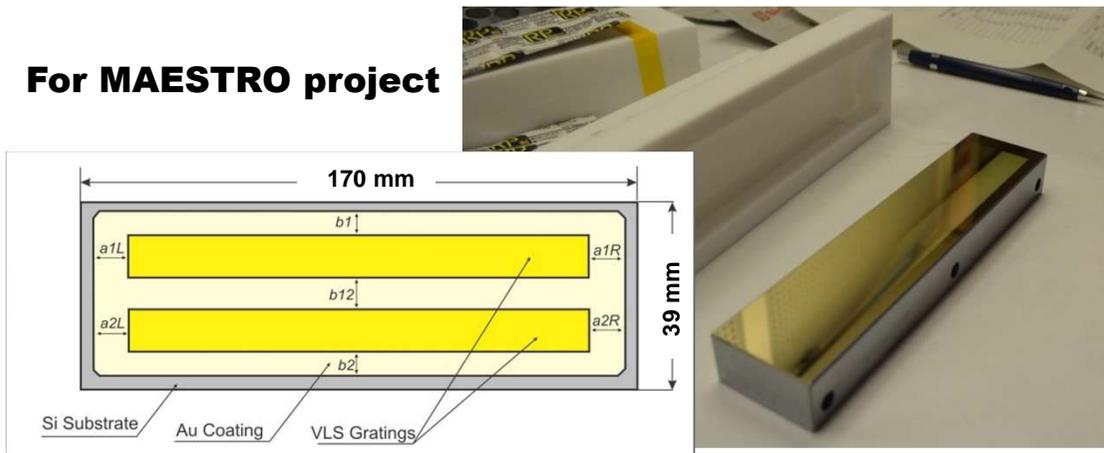
Characterization of VLS gratings with ZYGO NewView™-7300 interferometric microscope



MEASURING CAPABILITIES:

- Built-in PSD processing of the measured surface height distributions;
- Broad range of spatial resolution with different objectives (2.5x – 50x) and zooms (0.5x, 1x, and 2x);
- Multiple repeatable measurements with automated incremental translation of surface under test;
- Measurements with stitching...

For MAESTRO project



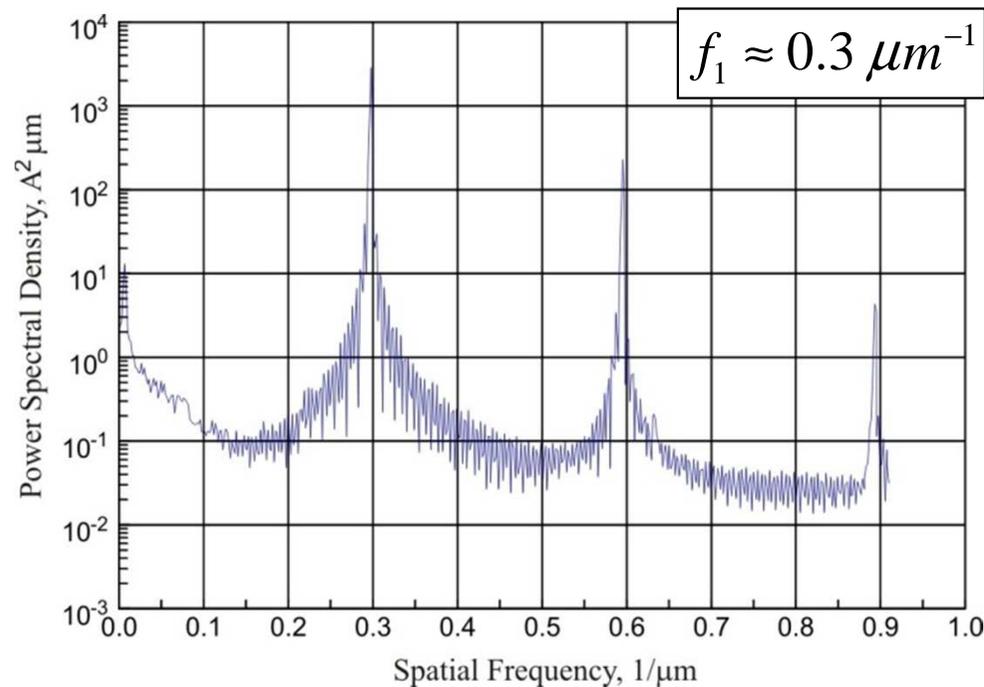
PSD spectrum of a 300-line-per-mm grating

Measured with the ZYGO NewView™-7300
interferometric microscope ($\lambda = 0.57 \text{ nm}$)
using **20×** objective and the **1×** zoom:

effective pixel size: $\Delta p \approx 0.11 \text{ } \mu\text{m}$

Nyquist frequency: $F_N \approx 0.91 \text{ } \mu\text{m}^{-1}$

number of PSD peaks: $n = 8$



PSD spectrum of the grating consists of n
peaks at the spatial frequencies:

$$\omega_n = 2\pi n / D = 2\pi f_n,$$

$$1 \leq n < F_N / f_1,$$

where, the Nyquist frequency of the microscope:

$$F_N = 1 / 2\Delta p,$$

Δp is the effective pixel size of the CCD
detector.

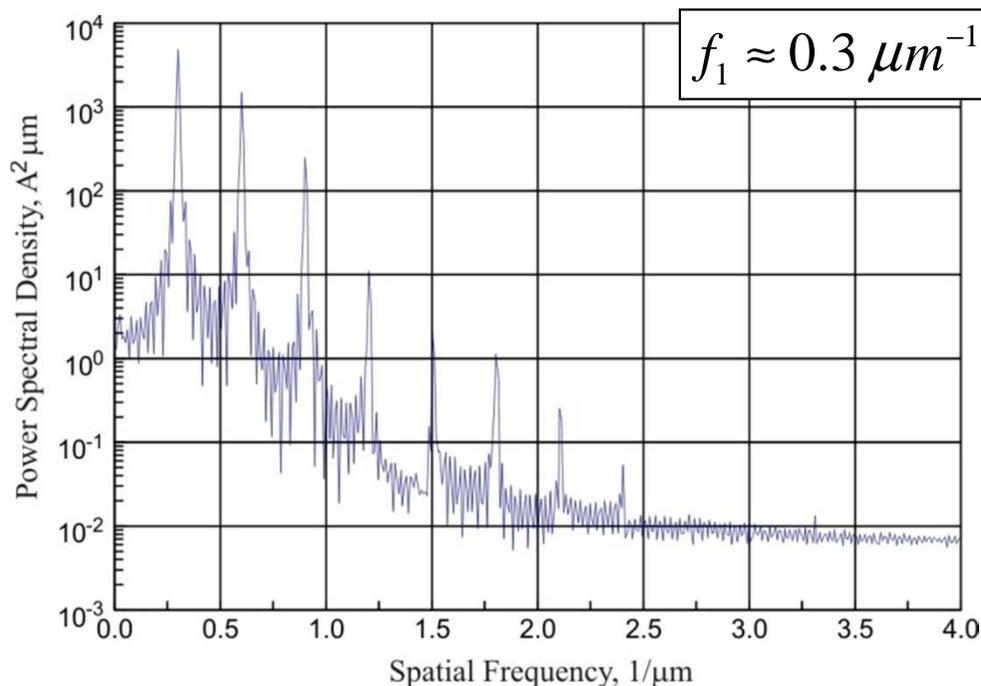
PSD spectrum of a 300-line-per-mm grating

Measured with the ZYGO NewView™-7300 interferometric microscope ($\lambda = 0.57 \text{ nm}$) using **50×** objective and the **2×** zoom:

effective pixel size: $\Delta p \approx 0.11 \text{ } \mu\text{m}$

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N.B. Microscope resolution is about

0.52 μm

determined by the diffraction limit of the objective:*

$$MTF_o(f) = \frac{2}{\pi} \left[-\Omega \sqrt{1 - \Omega^2} + \text{ArcCos } \Omega \right]$$

where, $\Omega = \lambda f / 2NA$, $f = \sqrt{f_x^2 + f_y^2}$,

NA is the Numerical Aperture.

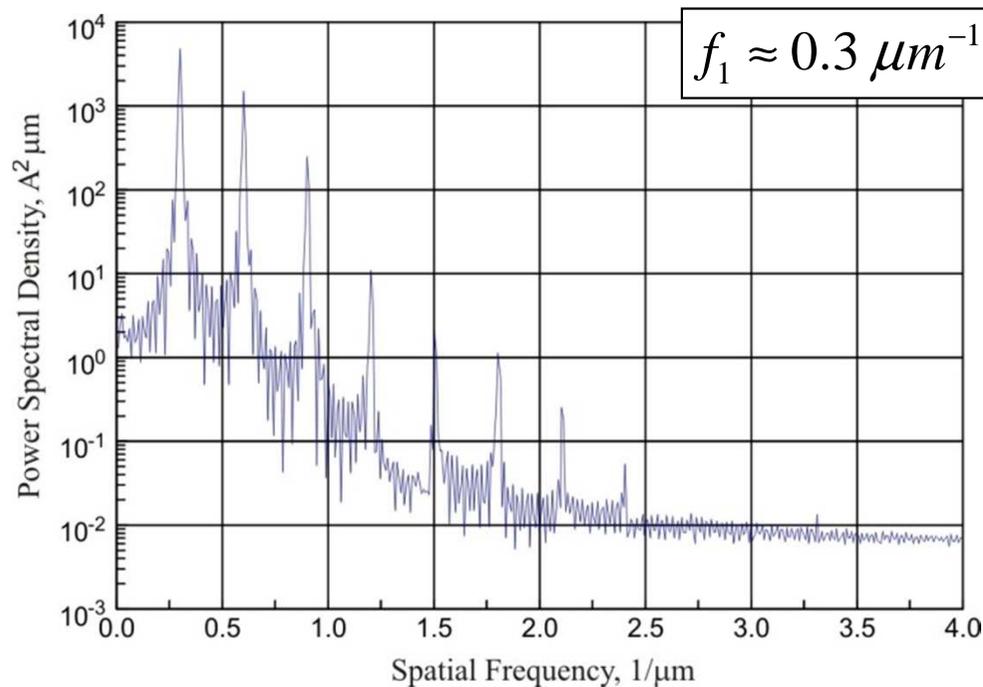
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QUESTIONS:

1. How to reliably position the PSD peaks?

- Analytical fitting function...
- Fitting procedure...

2. What is the best arrangement for the measurements?

- Microscope resolution (objective, zoom)...
- Measurements with stitching...

3. What is the reliability of the measurements?

- Spatial calibration...
- Cross-comparison with other methods...

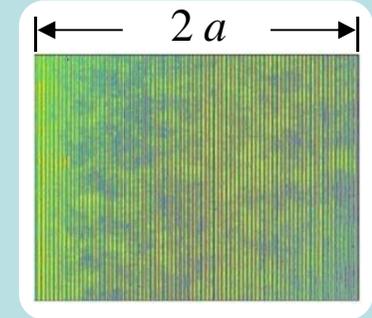
Line shape measured with a microscope with a finite size of field-of-view

If $[-a, +a]$ determines the size of the field-of-view in the direction along the grating:

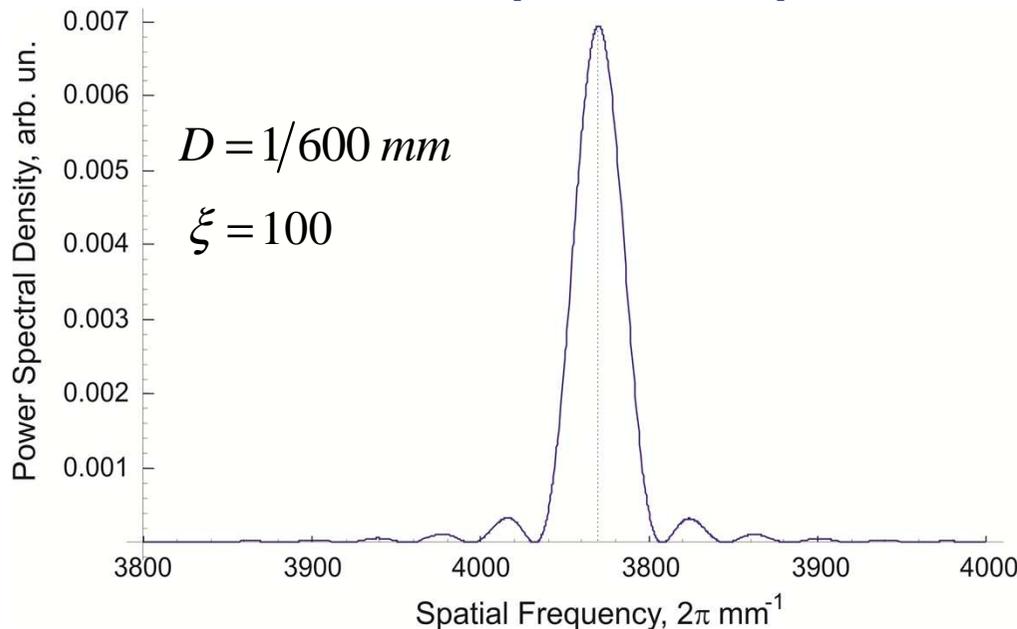
$$P_{f_1}(f) \propto I_1(f) = \left| \int_{-a}^a \text{Sin}\left(\frac{2\pi}{D}x\right) \exp(-ifx) dx \right|^2$$

$$= 4 \left[f_1 \text{Cos}(af_1) \text{Sin}(af) - f \text{Cos}(af) \text{Sin}(af_1) \right]^2 / (f^2 - f_1^2)^2$$

$$= 4 \left[\frac{2\pi}{D} \text{Cos}(\pi\xi) \text{Sin}\left(\frac{\xi D}{2}f\right) - f \text{Cos}\left(\frac{\xi D}{2}f\right) \text{Sin}(\pi\xi) \right]^2 / \left(f^2 - \left(\frac{2\pi}{D}\right)^2 \right)^2$$



First harmonic peak in PSD spectrum



$\xi = 2a/D$ is the number of grooves seen in the microscope's field-of-view

If ξ is an integer number:

$$I_1(f) = 4 \left[\frac{2\pi}{D} \text{Sin}\left(\frac{\xi D}{2}f\right) \right]^2 / \left(f^2 - \left(\frac{2\pi}{D}\right)^2 \right)^2$$

$$= \left\{ 2\pi\xi \cdot \text{Sinc}\left(\frac{\xi D}{2}\left(f - \frac{2\pi}{D}\right)\right) / \left(f + \frac{2\pi}{D}\right) \right\}^2$$

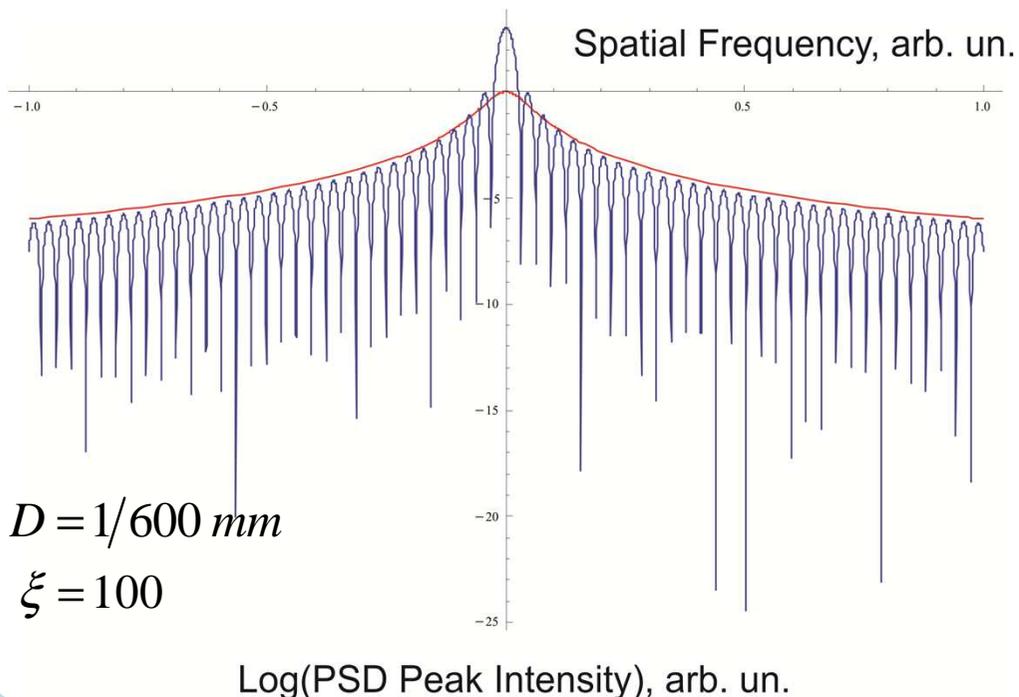
Logarithmic scale to increase the number of points included into the peak fitting

$$I_1(f) = \left\{ 2\pi\xi \cdot \text{Sinc} \left(\frac{\xi D}{2} \left(f - \frac{2\pi}{D} \right) \right) \right\}^2$$

$$S(f) = \text{Log} [I_1(f)] = A_0 + B_0 \cdot \text{Log} \left| \text{Sinc} \left(\frac{\xi D}{2} \left(f - \frac{2\pi}{D} \right) \right) \right| + C_0 \cdot \text{Log} \left(f + \frac{2\pi}{D} \right)$$

$$= A_0 + B_0 \cdot \text{Log} \left| \text{Sinc} (a \cdot (f - f_1)) \right| + C_2 \cdot (f - f_1)$$

First harmonic peak in PSD spectrum



Problems to solve:

- Fit in the presence of the oscillations
- Fit at the peak center

Fit with the Lorentzian function:

$$L(f) \approx A_L + B_L \text{Log} \left[\frac{(w/2)^2}{f^2 + (w/2)^2} \right]$$

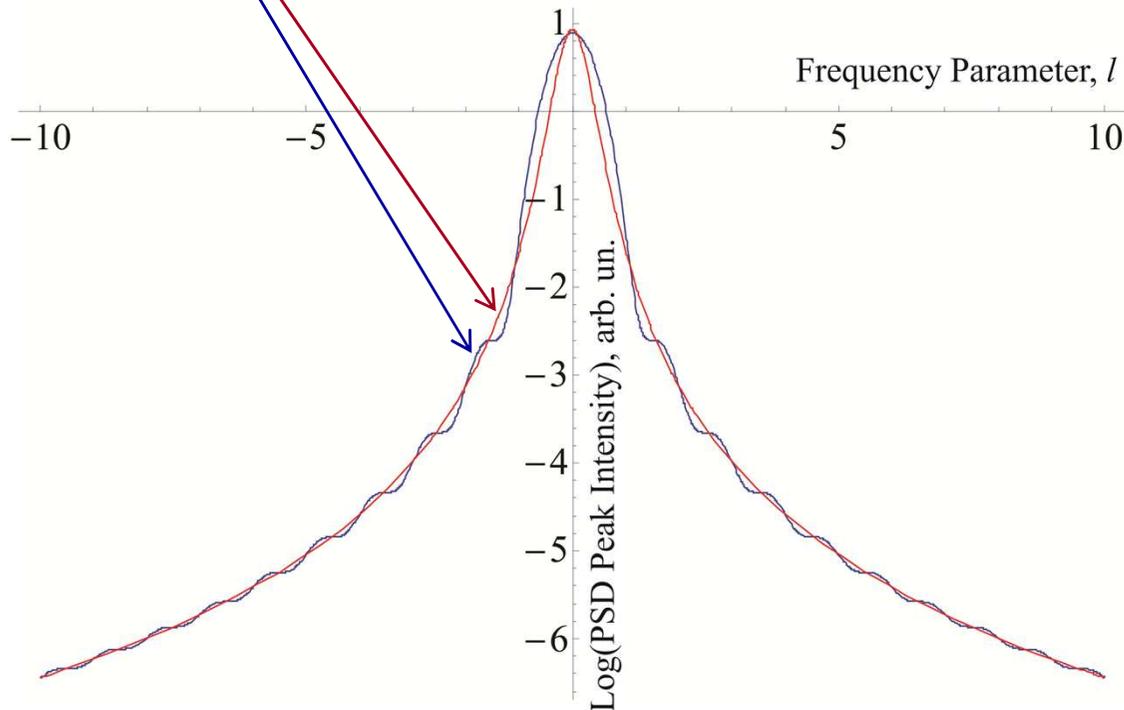
Smoothing of *Sinc*-function oscillations in discrete PSD measurements

Averaging over half of the oscillation period: $\Delta f = 2\pi/\xi D$

$$S_{Int}(f_l) = \int_{\pi l - \pi/2}^{\pi l + \pi/2} \text{Sinc}^2(x) dx$$

$$= \frac{4\text{Cos}^2(l\pi) + (4l^2 - 1)\pi [\text{SinIntegral}(\pi - 2l\pi) + \text{SinIntegral}(\pi + 2l\pi)]}{(4l^2 - 1)\pi}$$

$$\approx A + B \cdot \text{Log} \left[\frac{((f - f_1)^2 + \eta \cdot (w/2)^2)}{((f - f_1)^2 + (w/2)^2)^2} \right] + C \cdot (f - f_1)$$

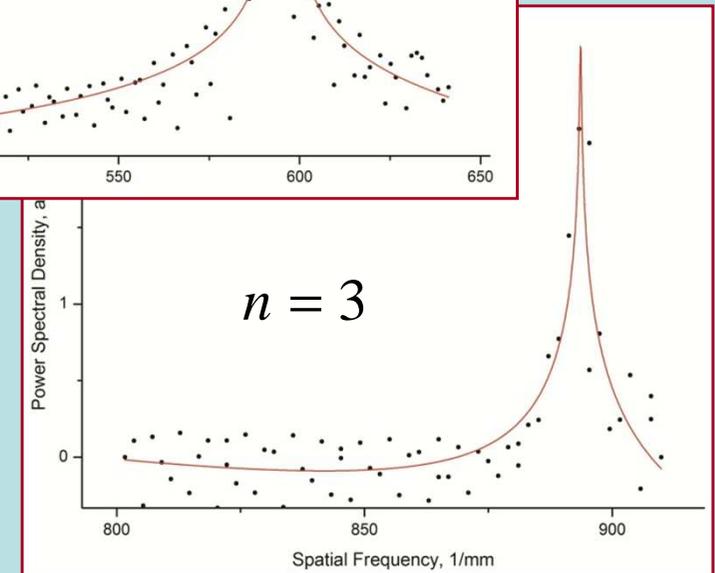
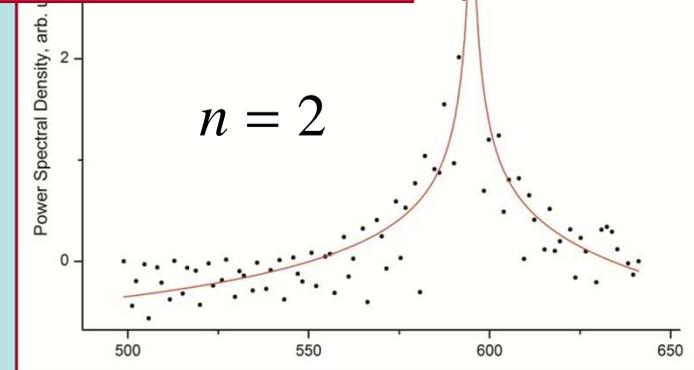
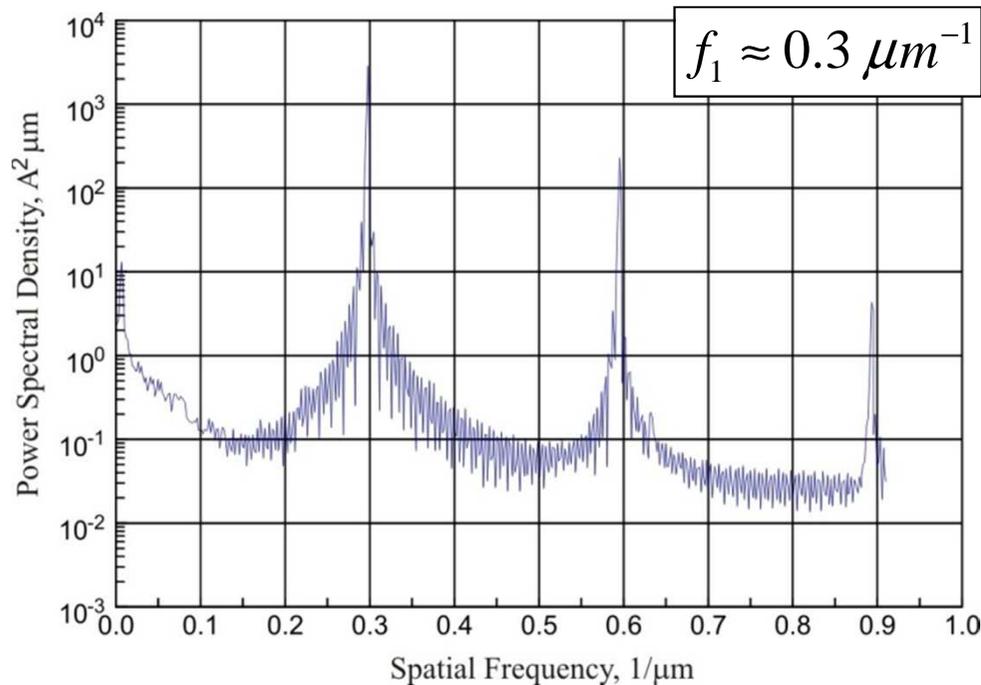
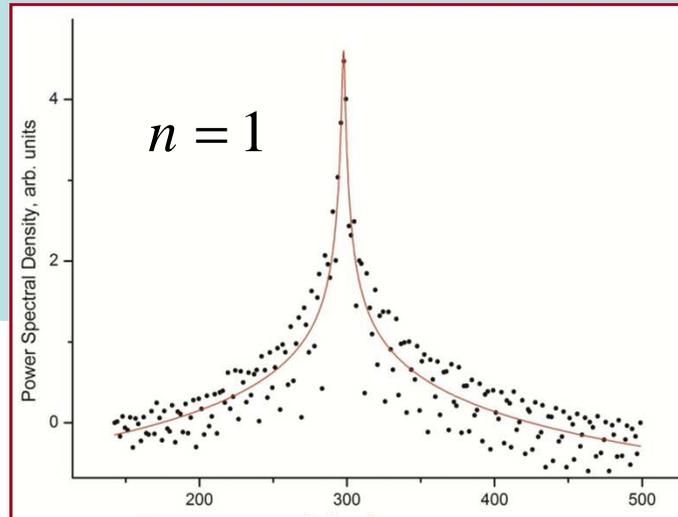


N. B. If the number of points per one grating period is just a few, the PSD measurements strongly depend on mutual alignment of the grating groove phase with respect to the pixel grid of the detector CCD camera.

This case should be specially analyzed (in progress).

Fitting of the PSD peaks with the derived analytical function

Measured with the ZYGO NewView™-7300 interferometric microscope ($\lambda = 0.57 \text{ nm}$) using $20\times$ objective and the $1\times$ zoom:



$$S_{Int}(f_i) = A + B \cdot \text{Log} \left[\frac{(f - f_1)^2 + \eta \cdot (w/2)^2}{(f - f_1)^2 + (w/2)^2} \right]^2 + C \cdot (f - f_1)$$

Evaluation of the groove density

from the entire PSD spectrum

ZYGO NewView™-7300 with the 300-line-per-mm grating

20× objective and the 1× zoom

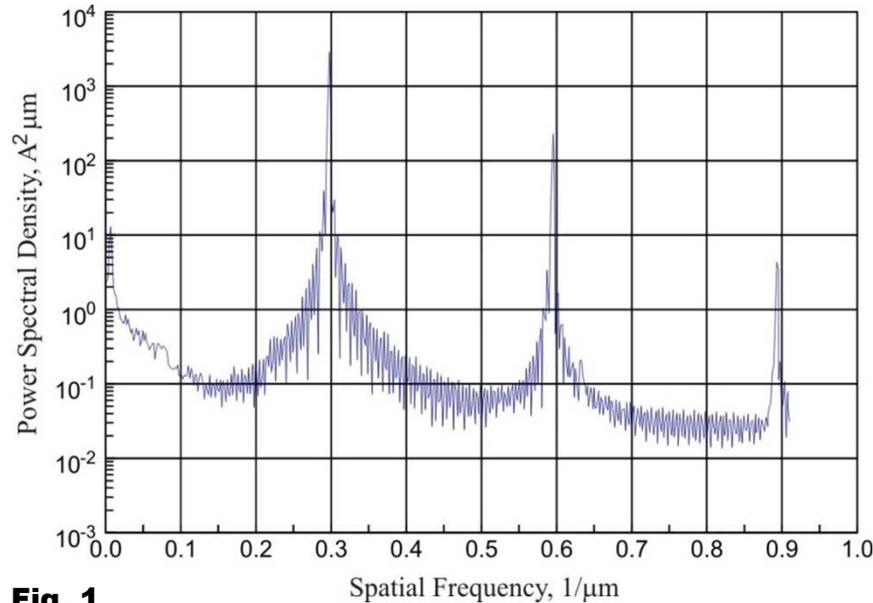


Fig. 1

50× objective and the 2× zoom

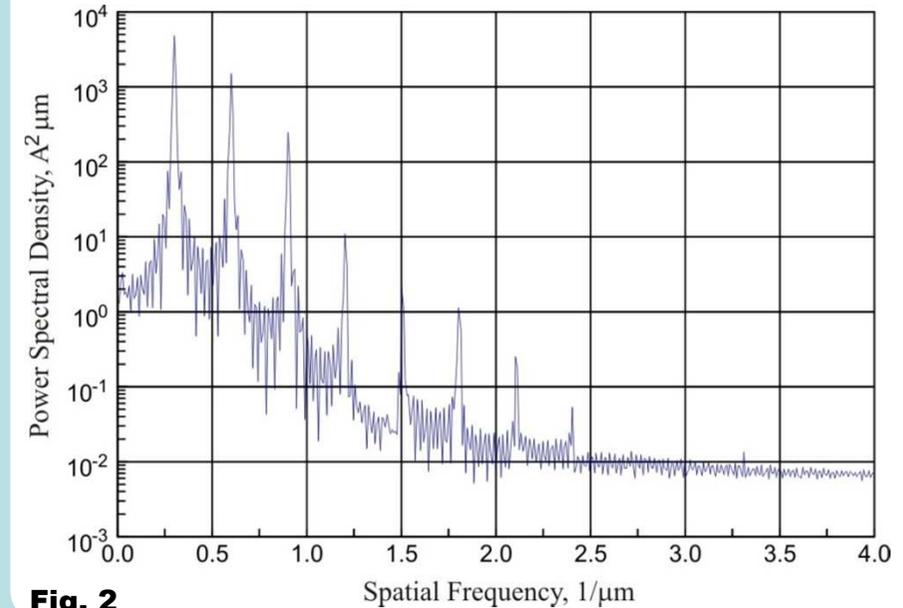


Fig. 2

n	PSD spectrum in Fig. 1		PSD spectrum in Fig. 2	
	Peak Frequency, 1/mm	Standard Error, 1/mm	Peak Frequency, 1/mm	Standard Error, 1/mm
1	297.53	0.30	302.68	1.9
2	595.01	0.40	603.80	2.3
3	893.60	0.64	905.09	1.8
4	--	--	1204.77	2.5
5	--	--	1507.83	2.2
6	--	--	1805.66	3.0
7	--	--	2104.72	4.5
8	--	--	2395.83	12.5

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Evaluation of the groove density from the entire PSD spectrum

ZYGO NewView™-7300 with the 300-line-per-mm grating

20× objective and the 1× zoom

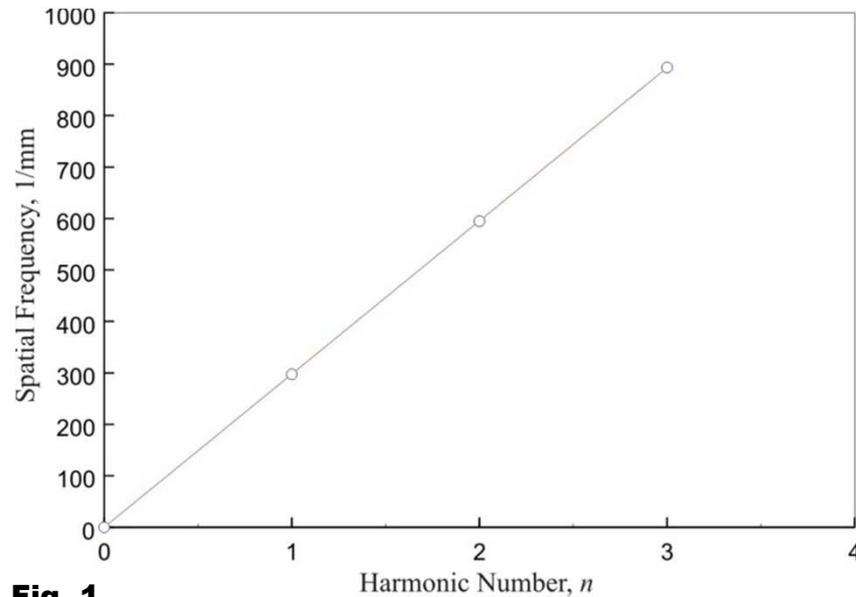


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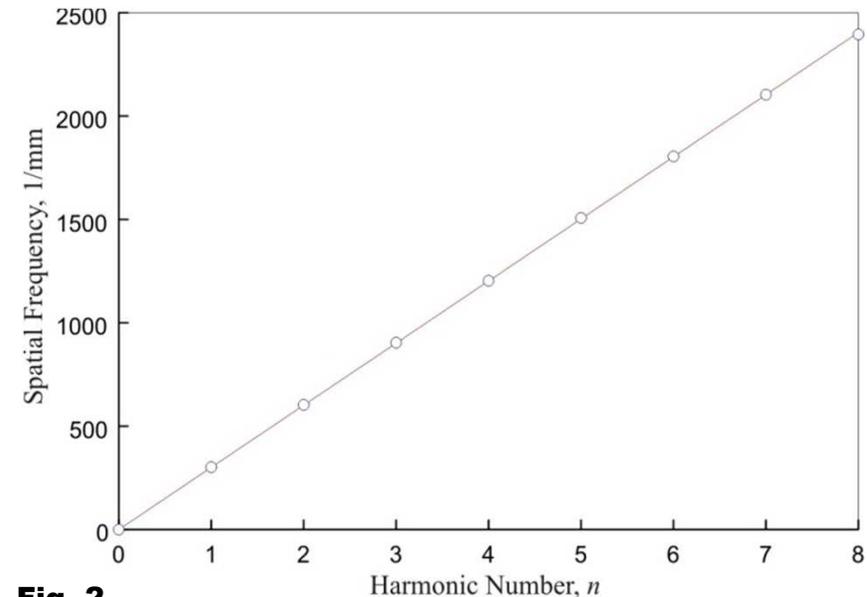


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Averaged Density	297.74 lines/mm	0.10 lines/mm	301.27 lines/mm	0.19 lines/mm

Number of grooves vs Number of peaks

ZYGO NewView™-7300 with the 300-line-per-mm grating

OBSERVATION:

The linear fit of data in Fig. 2 provides an improvement of the standard error by a factor of **~10**.

- **Fitting with weighting accounts for the standard errors of the peak frequency values;**
- **Statistical significance of a peak in the final value of the groove density is proportional to the harmonic number, n .**

The standard error of a peak frequency is inversely proportional to the linear size of the field of view.*

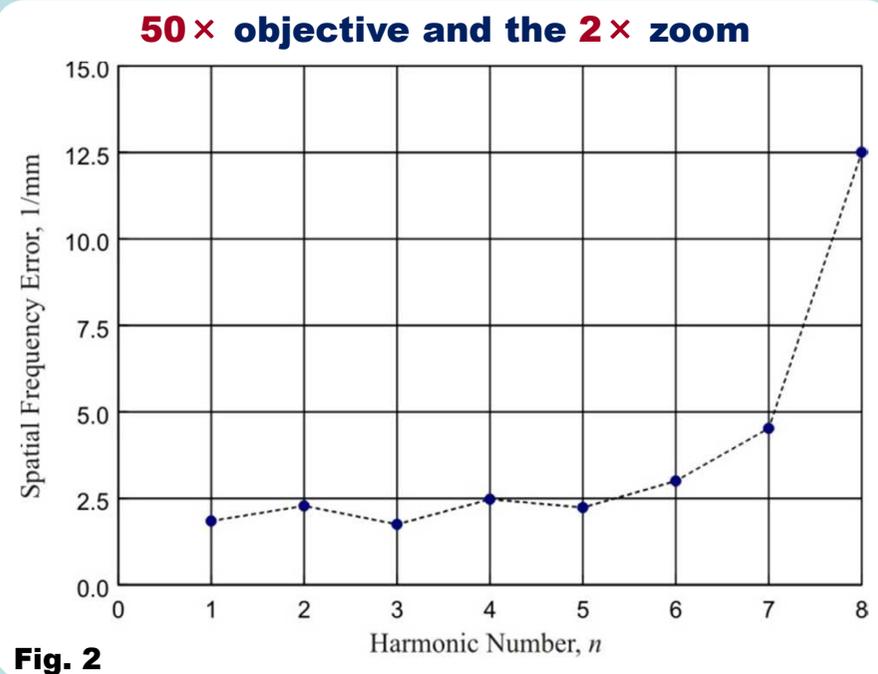


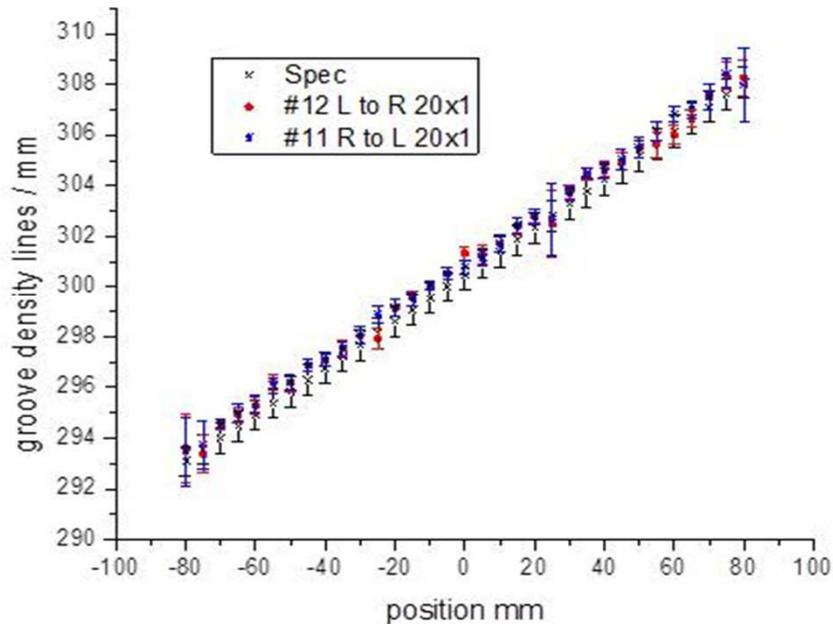
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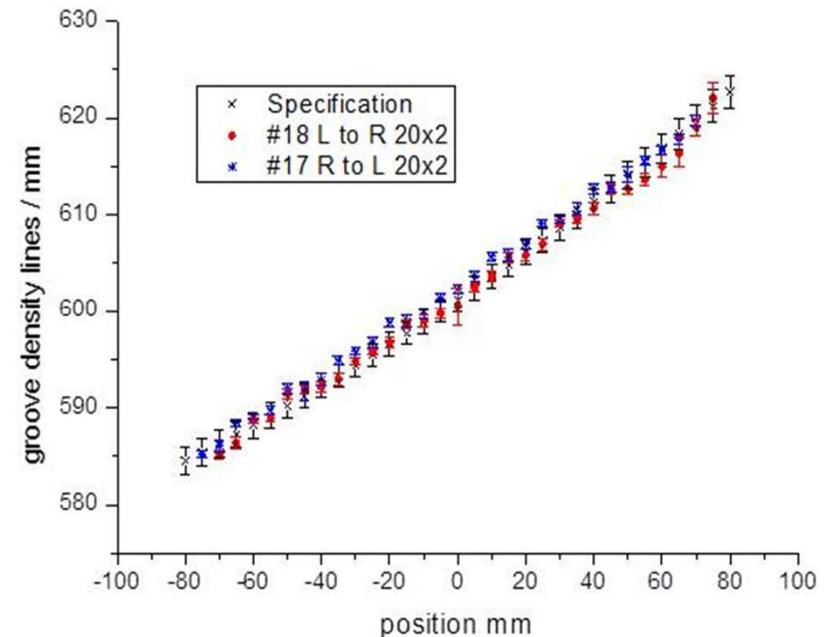
* D. L. Voronov, R. Cambie, R. M. Feshchenko, E. Gullikson, H. A. Padmore, A. V. Vinogradov, V. V. Yashchuk, *Development of an ultra-high resolution diffraction grating for soft X-rays*, Proc. of SPIE Vol. 6705, 67050E (2007).

Cross-comparison with vendor metrology

300-lines-per-mm grating



600-lines-per-mm grating



	Specification (mm)	Measurement (mm)	Specification (mm)	Measurement (mm)
g_0	$300 \pm 0.6 \text{ mm}^{-1}$	$300.87 \pm 0.05 \text{ mm}^{-1}$	$600 \pm 1.2 \text{ mm}^{-1}$	$602.02 \pm 0.11 \text{ mm}^{-1}$
g_1	$0.0935 \pm 0.00047 \text{ mm}^{-2}$	$0.0920 \pm 0.0008 \text{ mm}^{-2}$	$0.235 \pm 0.001 \text{ mm}^{-2}$	$0.23091 \pm 0.0018 \text{ mm}^{-2}$
g_2	$1.9 \times 10^{-5} \pm 0.29 \times 10^{-5} \text{ mm}^{-3}$	$2.08 \times 10^{-5} \pm 1.4 \times 10^{-5} \text{ mm}^{-3}$	$3.8 \times 10^{-5} \pm 0.57 \times 10^{-5} \text{ mm}^{-3}$	$9.07 \times 10^{-5} \pm 3.1 \times 10^{-5} \text{ mm}^{-3}$

"truth" is hard to know; it may also have uncertainty...

Results obtained:

- **An analytical expression for the shape of the first harmonic peak in the PSD spectrum of a diffraction grating has been derived.**
- **Based on the analytical expression, an approximate fitting function has been suggested.**
- **The developed fitting procedure has been successfully applied to the PSD metrology of a 300-lines-per-mm grating.**
- **Cross-comparison with the vendor metrology has suggested for high reliability of the lateral calibration of the interferometric microscope used for the measurements.**
- **A possibility to get valuable data on the groove density for a grating with a pitch smaller than the microscope resolution has been demonstrated.**

Questions to be investigated (research in progress):

- **Dependence of the PSD measurements on mutual alignment of the grating groove phase with respect to the pixel grid of the detector CCD camera in the case when the number of points per one grating period is just a few.**
- **Theory of groove density measurements for a grating with a pitch smaller than the microscope's resolution (e.g., via use of the aliasing effect).**
- **Ways to increase accuracy: Measurements with stitching... to test stitching!!!**
(Note that the first try has failed...)

We are grateful to Tony Warwick for providing the MAESTRO grating for optical metrology.

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“Measurement science is not, however, purely the preserve of scientists. It is something of vital importance to us all. The intricate but invisible network of services, suppliers, and communications upon which **we are all dependent relies on metrology for its efficient and reliable operation.**”



Bureau International des Poids et Mesures (BIPM)
[Eng.: International Bureau of Weights and Measures]
<http://www.bipm.org/en/convention/wmd/2004/>

THANK YOU!